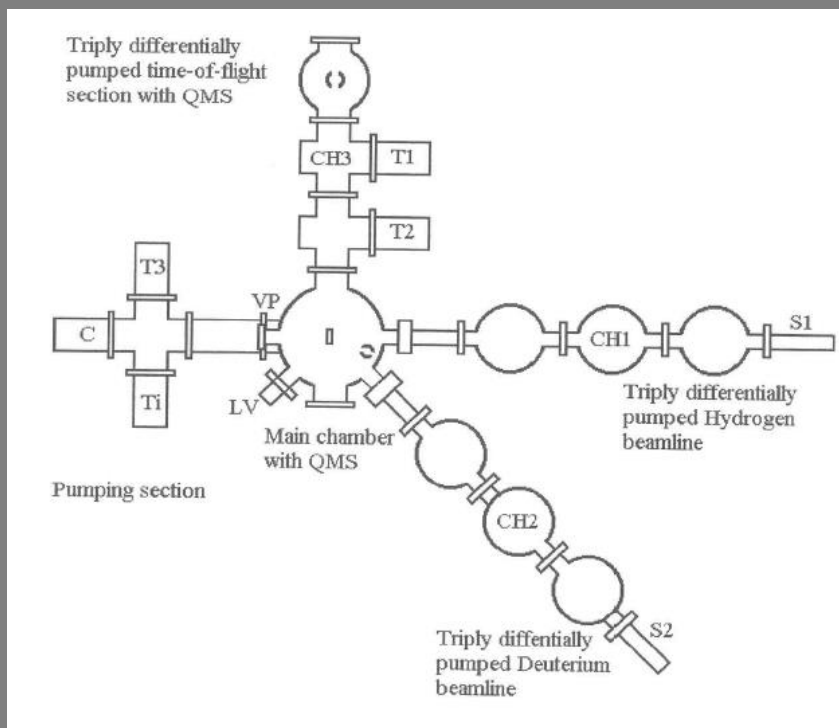


Mineral-biomolecules interactions: detect organic compounds in space



J.R. Brucato

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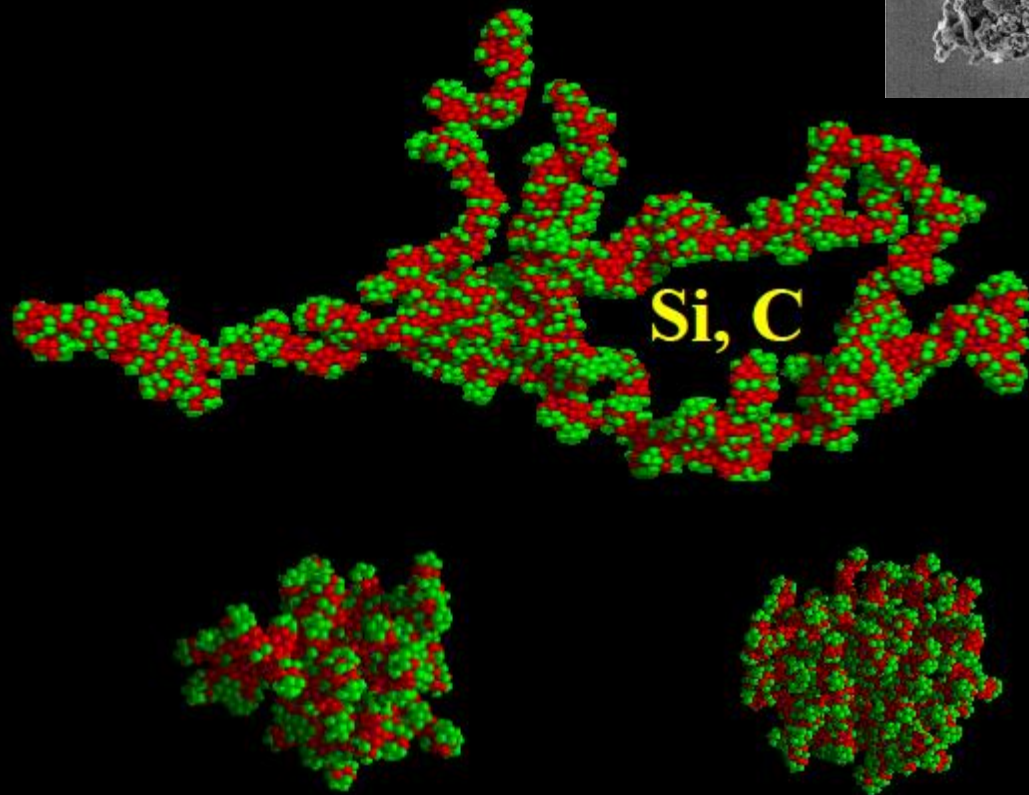
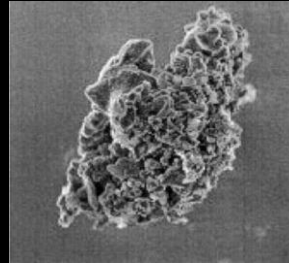


There are a number of space processes that can produce significant organic complex products:

- Gas phase ion-molecule interactions;
- Low-temperature gas-grain reactions;
- Gas phase unimolecular photodissociation;
- Ultraviolet photolysis and photodesorption in ice mantles
- Isotopic substitution due to ion irradiation.

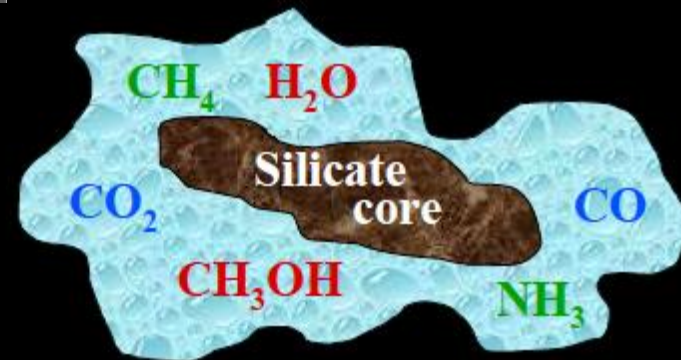
Dust particles: the seeds of planets and molecules

$T > 100 \text{ K}$



$T < 100 \text{ K}$

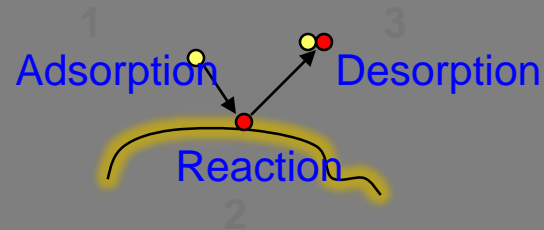
Ice coating



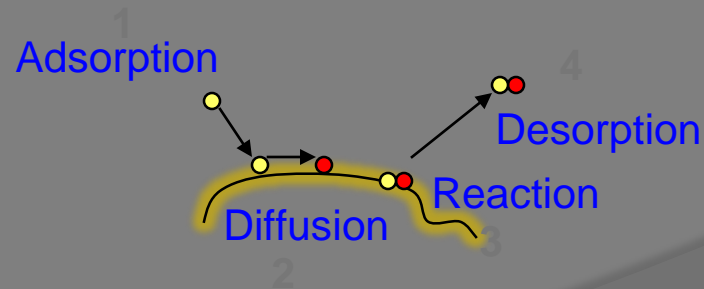
Surface catalysis

Surface catalysis allow molecules formation that are not possible in the gas phase. It opens pathways for the chemical evolution in space.

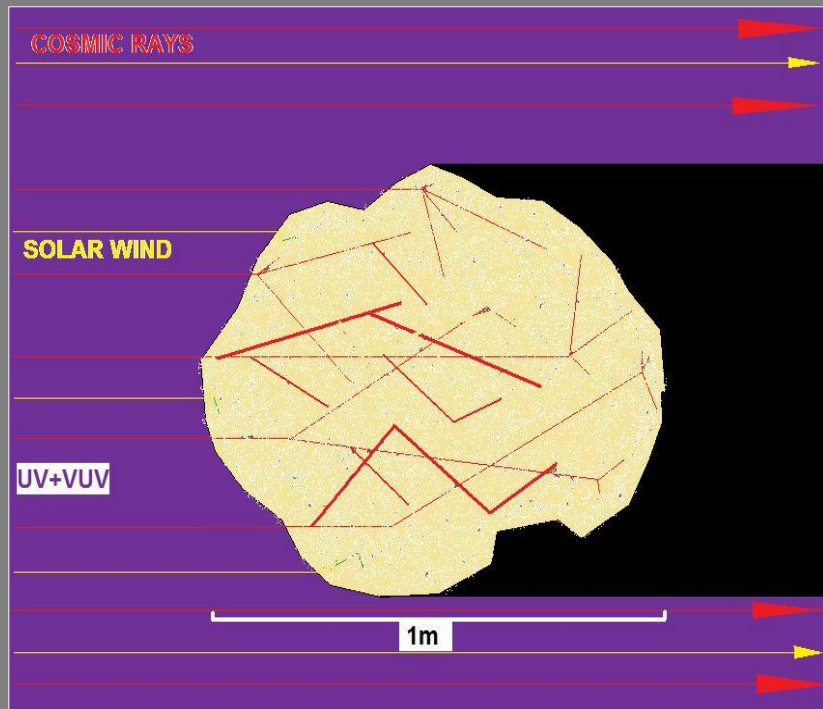
Eley-Rideal mechanism



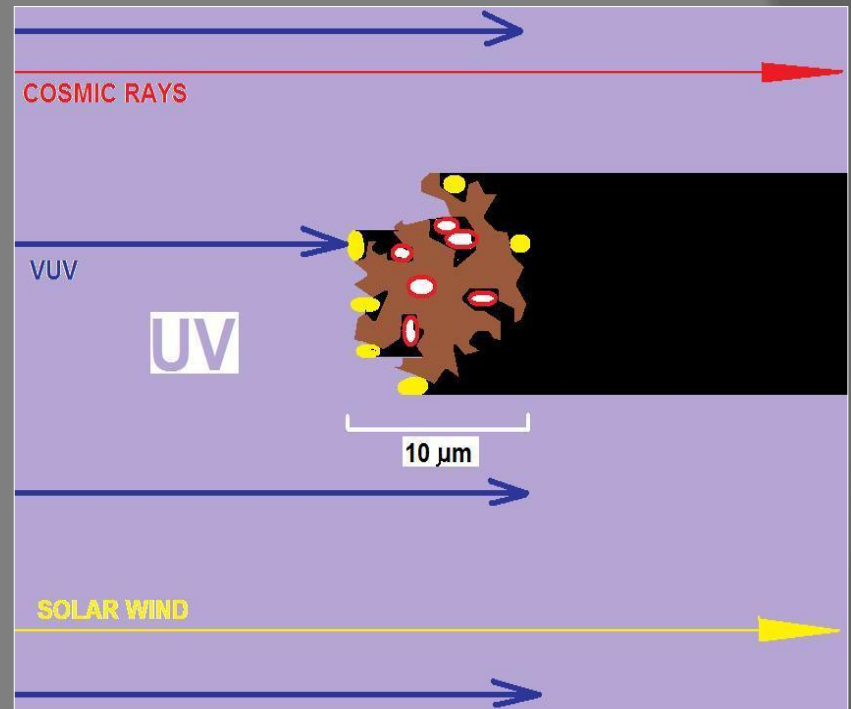
Langmuir-Hinshelwood mechanism



Big Rocks

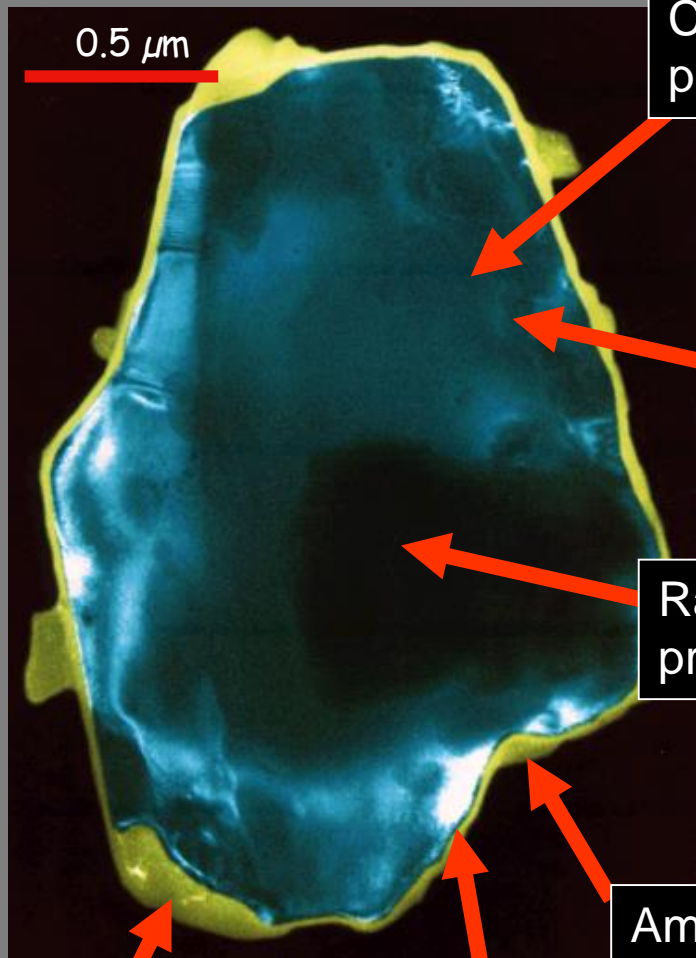


Small grains



Energy inputs (%) and Fluxes ($\text{cm}^{-2} \text{s}^{-1}$)

Solar Photons	2 eV	Visible (50%)	$2.0 \cdot 10^{17}$
	4 eV	NUV (10%)	$1.5 \cdot 10^{16}$
	6 eV	FUV (0.02%)	$3.0 \cdot 10^{13}$
Solar Wind (1 AU)	1keV	H ⁺ (95%)	$3.0 \cdot 10^8$
	4keV	He ²⁺ (5%)	
Solar Flares (1 AU)	>1 MeV	H ⁺ (95%)	$10^{10} (\text{cm}^{-2} \text{yr}^{-1})$
	>1 MeV	He ²⁺ (5%)	
Galactic cosmic rays	>1 MeV	H ⁺ (87%)	10^1
	>1 MeV	He ²⁺ (12%)	



Cosmic-ray
produced

Indigenous
Volatiles (?)

Radiogenic
production

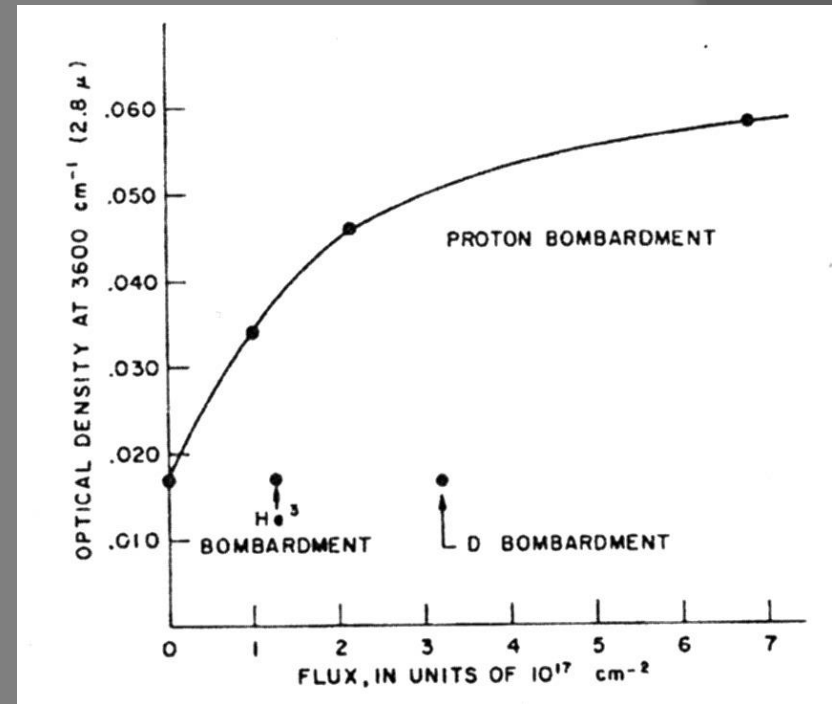
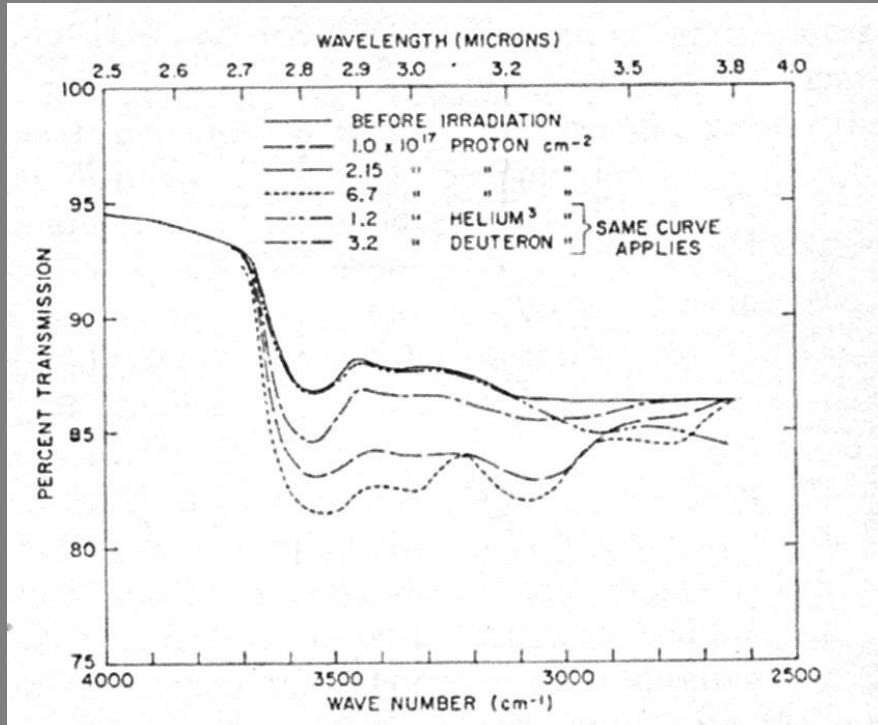
Amorphized skin

Solar wind

Non SW, surface-correlated

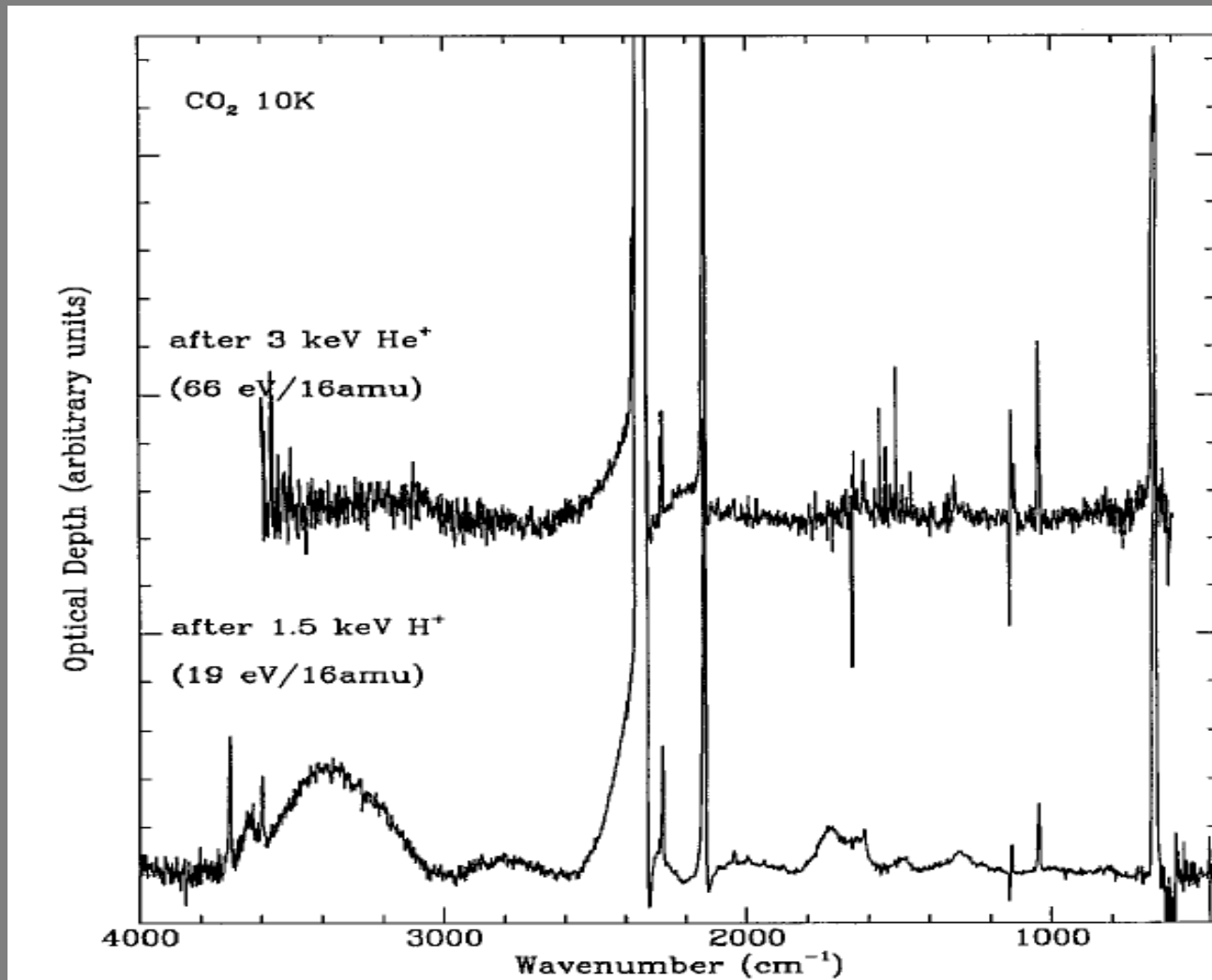


Proton-induced hydroxyl formation

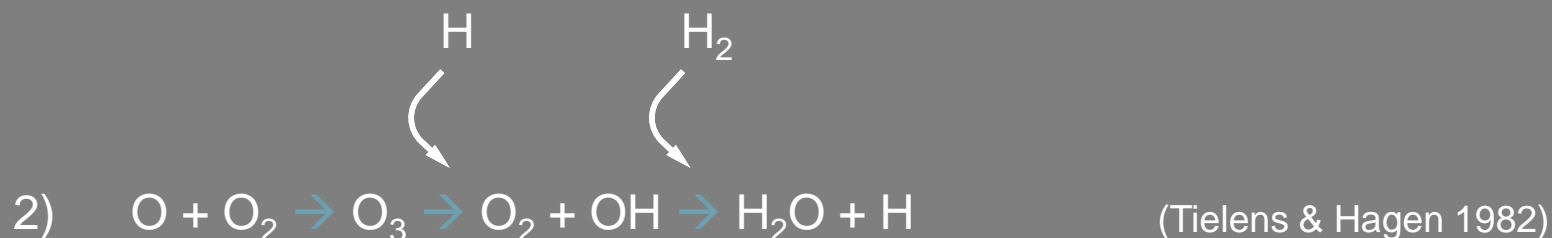
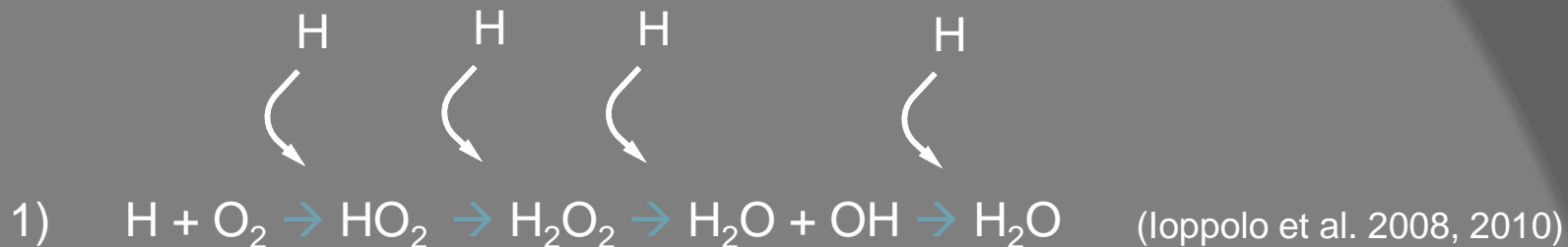


Estimate of the OH produced in the upper centimeter of the lunar surface material by proton bombardment: $4 \times 10^{16} \text{ OH cm}^{-3}$
(Zeller et al. 1966)

Water formation by ice irradiation



Brucato et al. 1997

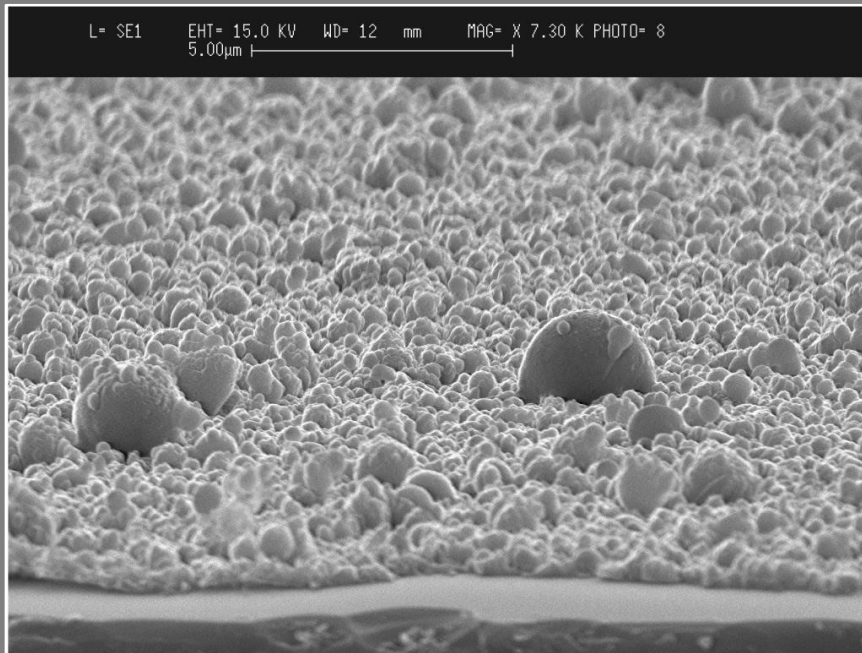


Some argue that species formed by these exothermic reactions will immediately desorb (Paoupluar 2005). However, models predict that most (99.1%) of OH and H₂O formed remain on surfaces (Cupper and Herbst 2007).

Silicate production in laboratory

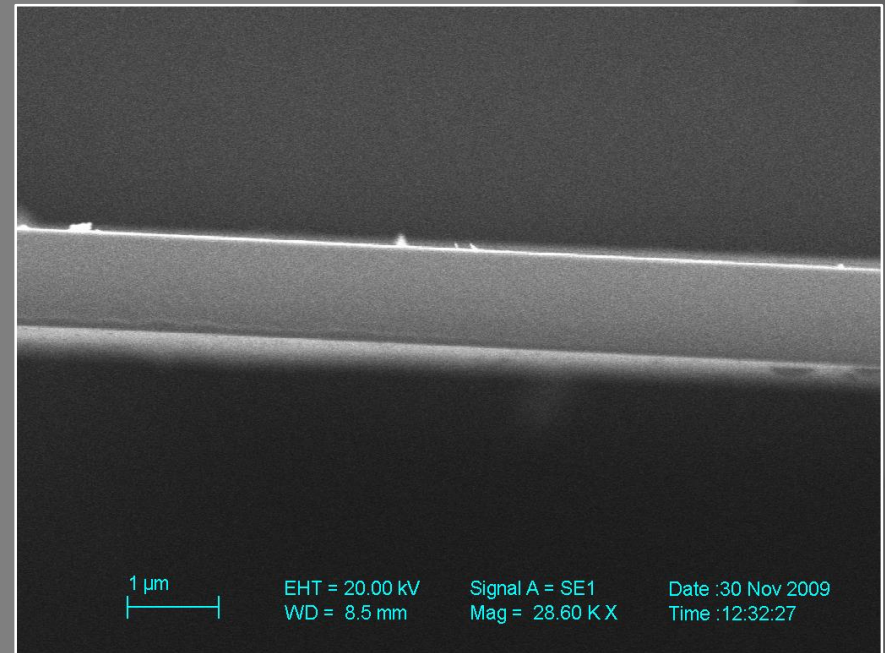
Amorphous olivine & pyroxene

Laser ablation

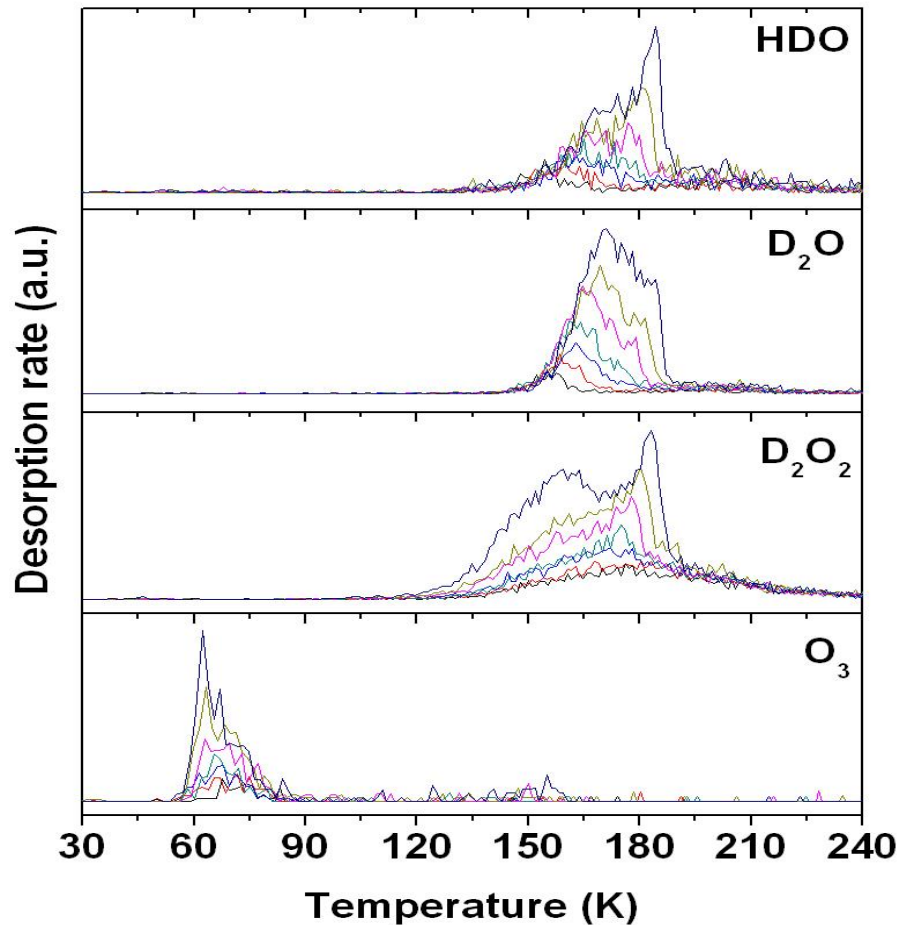


P= 10 mbar O₂

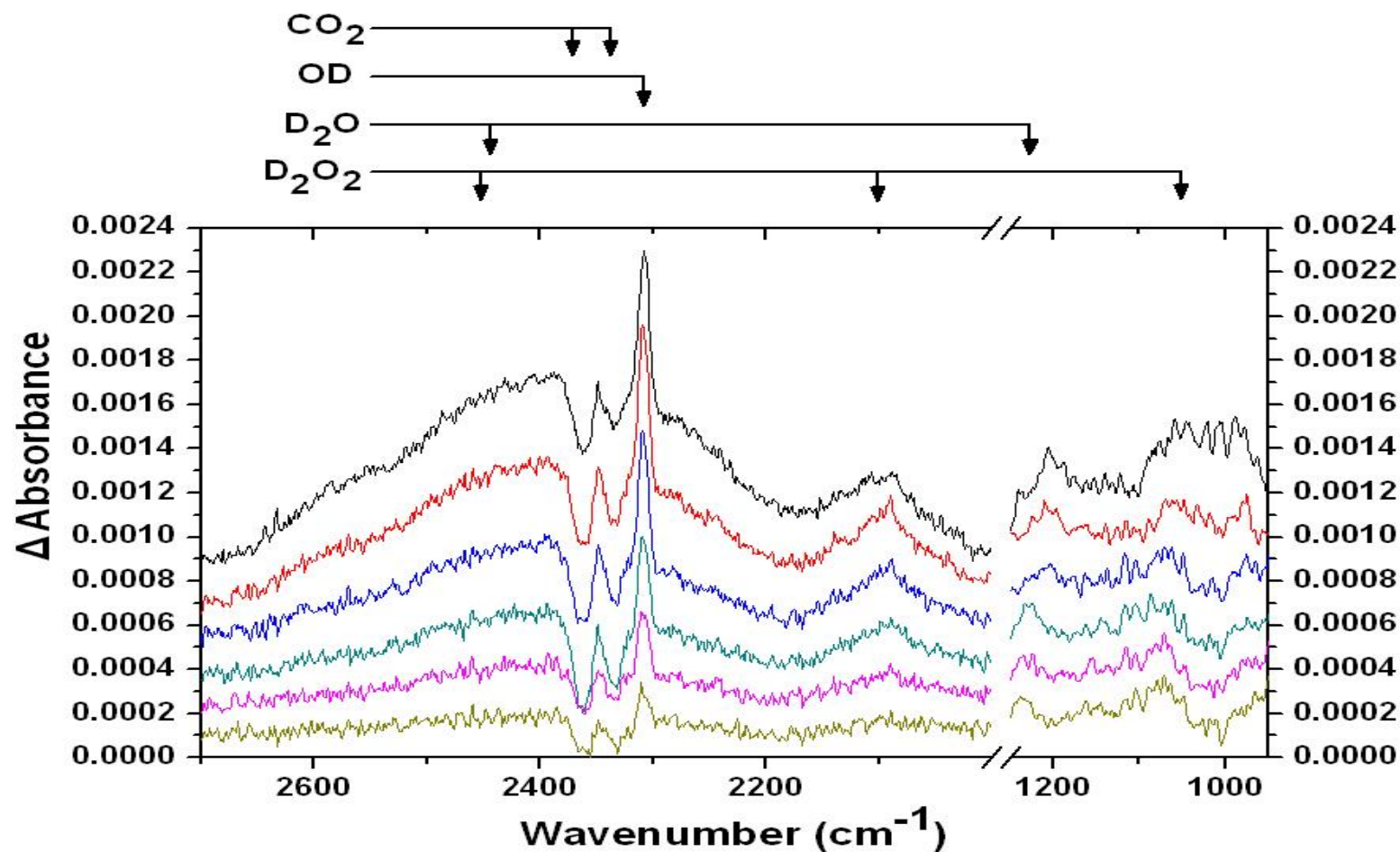
Electron Beam



P= 10⁻⁵ mbar

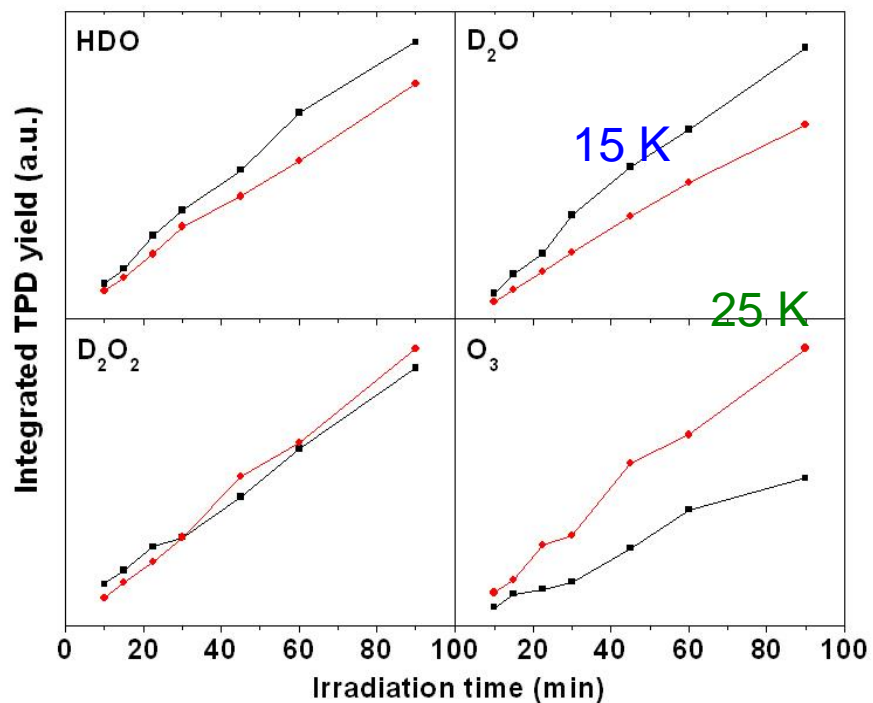


Desorption peaks for various species after D and O co-exposure. From bottom to top: 10 min, 15 min, 22.5 min, 30 min, 45 min, 60 min and 90 min.



RAIR spectra of D and O co-exposure for (from bottom to top) 1 hour, 2 hours, 3 hours, 4 hours, 5 hours and after annealing the sample at 70 K for 5 min after exposure. The spectra are displaced on the vertical axis for clarity.

Jing et al. 2011



Binding energy

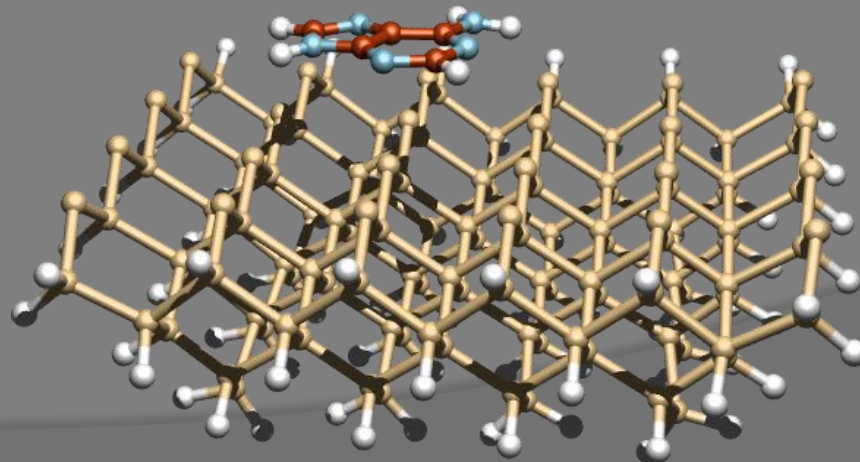
HDO 390 meV
D₂O 400 meV
H₂O 170 meV
D₂O₂ 430 meV

	HDO	D ₂ O	D ₂ O ₂	O ₃
15 K Slope (cm ⁻² min ⁻¹)	1.5×10 ¹³	4.2×10 ¹³	7.6×10 ¹³	2.6×10 ¹²
25 K Slope (cm ⁻² min ⁻¹)	1.2×10 ¹³	3.0×10 ¹³	8.9×10 ¹³	4.9×10 ¹²
15 K Formation efficiency	0.043	0.12	0.22	0.007
25 K Formation efficiency	0.034	0.087	0.26	0.014

Crucial transition from inanimate matter to biological systems probably occurred through **selection**, **concentration** and **organization of organic precursors**, yielding to the essential macromolecules of life



? MINERALS ?



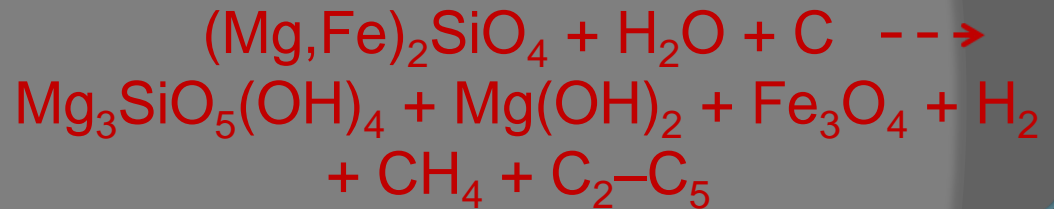
Interaction of Nucleic Acid Components with Mg-containing Minerals in two different Astrobiologically Relevant Environments:

1. Serpentine-hosted Hydrothermal Vents

Lost City Hydrothermal Field



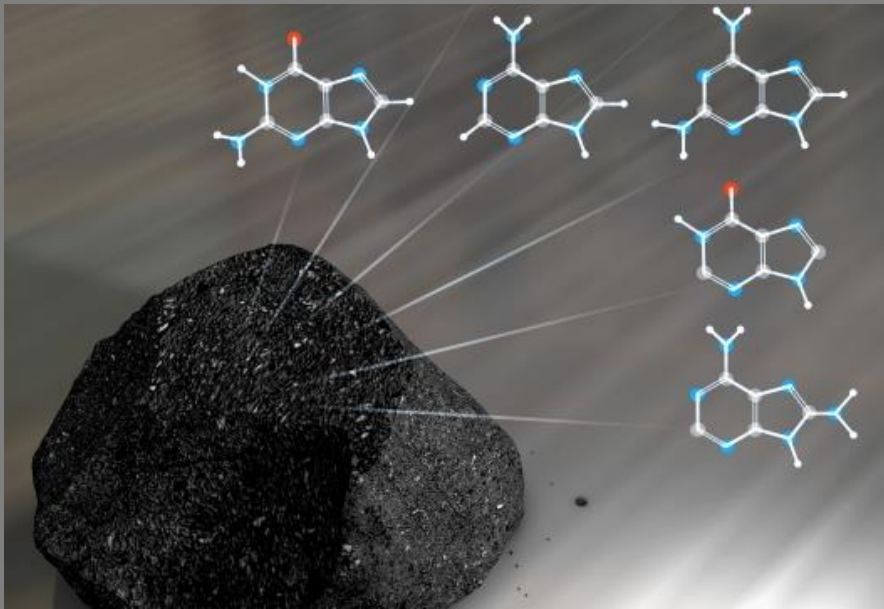
SERPENTINIZATION



- ✓ Disequilibria, redox gradient potentially catalyze formation of prebiotic molecules
- ✓ Lower temperatures typical of the Lost City hydrothermal fluids favor biosynthesis

Interaction of Nucleic Acid Components with Mg-containing Minerals in two different Astrobiologically Relevant Environments:

2. High UV Irradiation and Space-like Environments



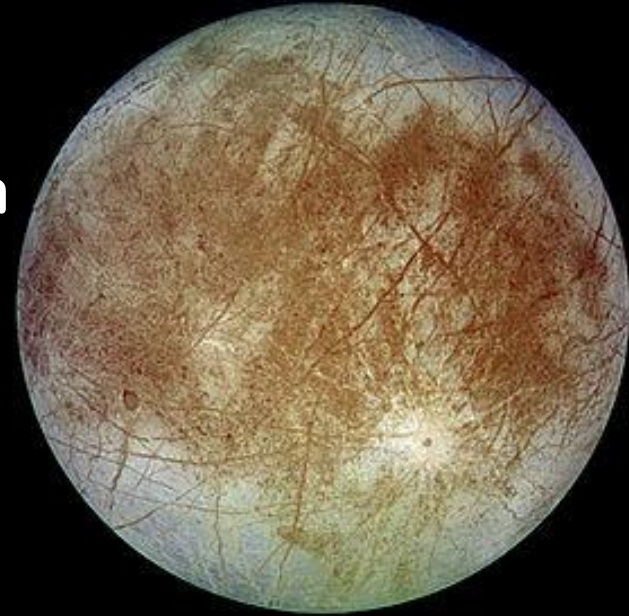
Magnesium oxide (MgO);
Forsterite (Mg₂SiO₄)

...

Weinbruch et al. *Meteoritics & Planetary Science* **2000**; Barber and Scott *PNAS* **2002**; Messenger et al. *Science* **2005**; Ming et al. *Journal of Geophysical Research* **2006**; Chevrier and Mathé *Planetary and Space Science* **2007**; Poteet et al. *The Astrophysical Journal Letters* **2011**.

POSSIBLE PRESENCE OF HYDROTHERMAL ACTIVITY

**Jupiter's moon
Europa**



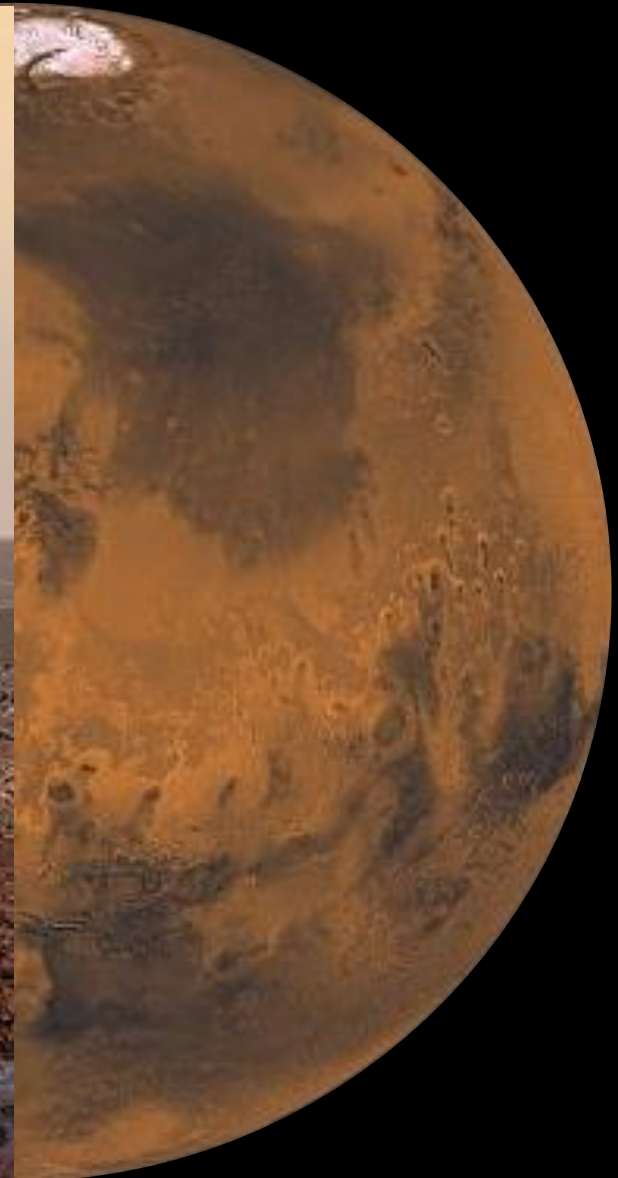
**Saturn's moon
Enceladus**



**Past Mars
(maybe
Noachian era ~
4.1 to 3.7 Gyr)**




ExoMars 2020



Astrobiology



Depth
exploring
ExoMars
rover

Mary Ann Liebert, Inc.  publishers

www.liebertpub.com/ast

Mars Organic Molecule Analyzer (MOMA)

ExoMars 2020

O1. Detect and characterize organic molecular species in solid samples with high (ppb) sensitivity.

O2. Analyze patterns of, and interrelationships among, a variety of organic molecules over a wide range of molecular weights and volatilities.

O3. Analyze for the presence and degree of ee in detected chiral organics.

O4. Characterize the inorganic geochemical context of the organic analyses.

R1. MOMA shall be able to detect organic molecules at concentrations as low as 10 ppb.

R2. MOMA shall be able to identify organics with chainbased or ring-based structures.

R3. MOMA shall be able to characterize molecular weight distribution patterns in organic molecules over mass-to-charge ratio as high as 1000 u.

R4. MOMA shall be able to detect compounds of low stability (e.g. , formaldehyde)

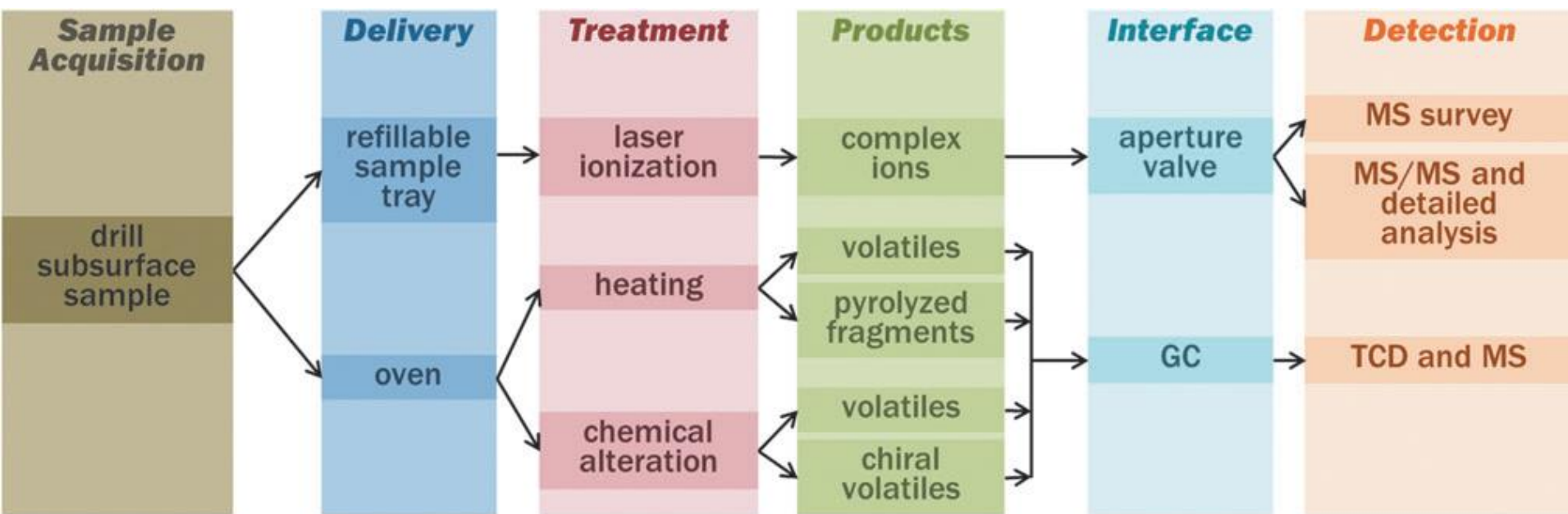
R5. MOMA shall investigate the biotic or abiotic origin of organic chiral molecules by analyzing their enantiomers at molecular concentrations as low as 1 ppmw.

R6. MOMA shall be able to detect light volatile organics at concentrations as low as 1 ppmw.

R7. MOMA shall be able to detect refractory organics (e.g. , heavy PAHs, kerogen-like material) with molecular weights up to 1000 u at concentrations as low as 10 ppmw.

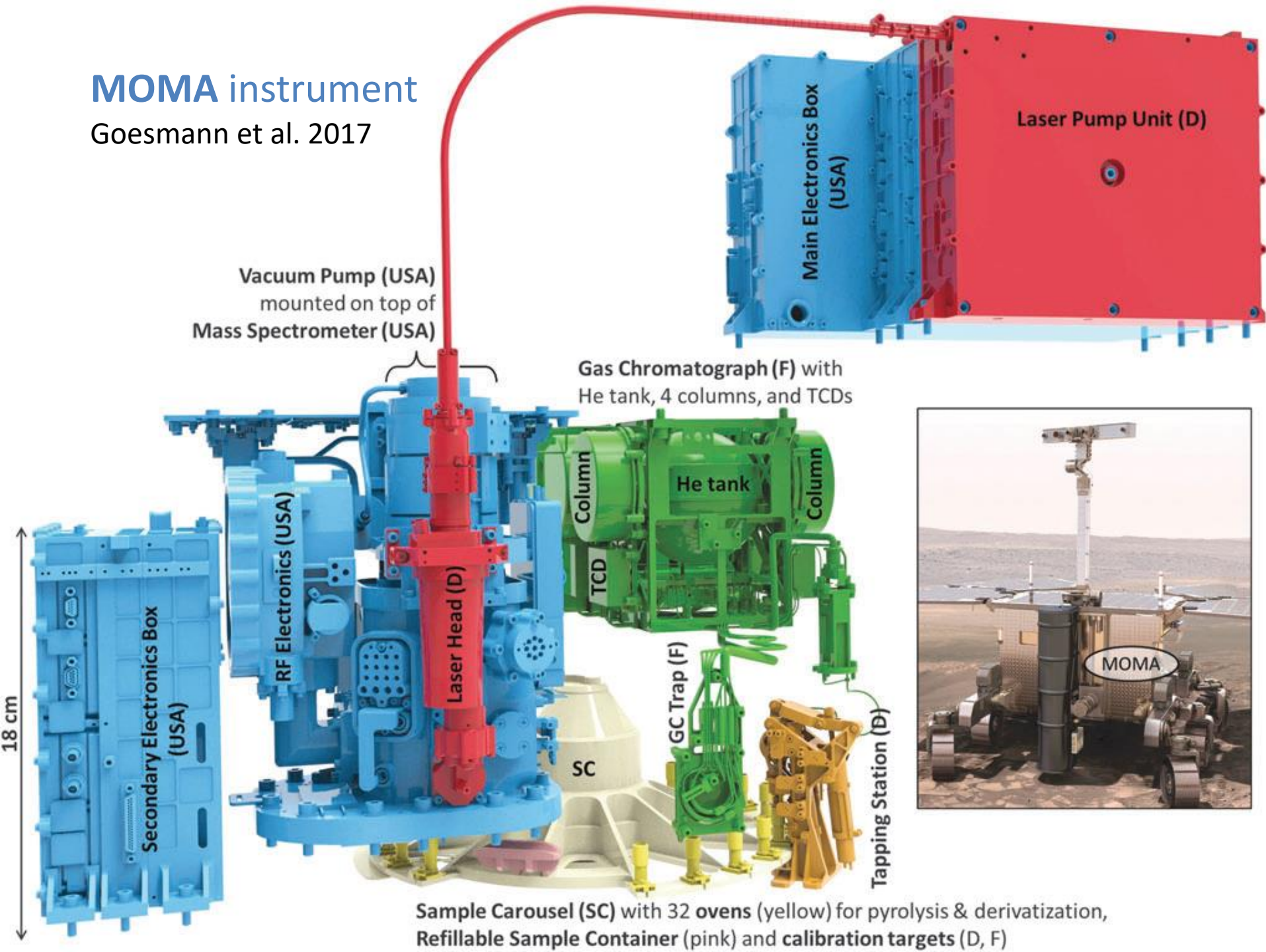
Mars Organic Molecule Analyzer (MOMA)

ExoMars 2020



MOMA instrument

Goesmann et al. 2017

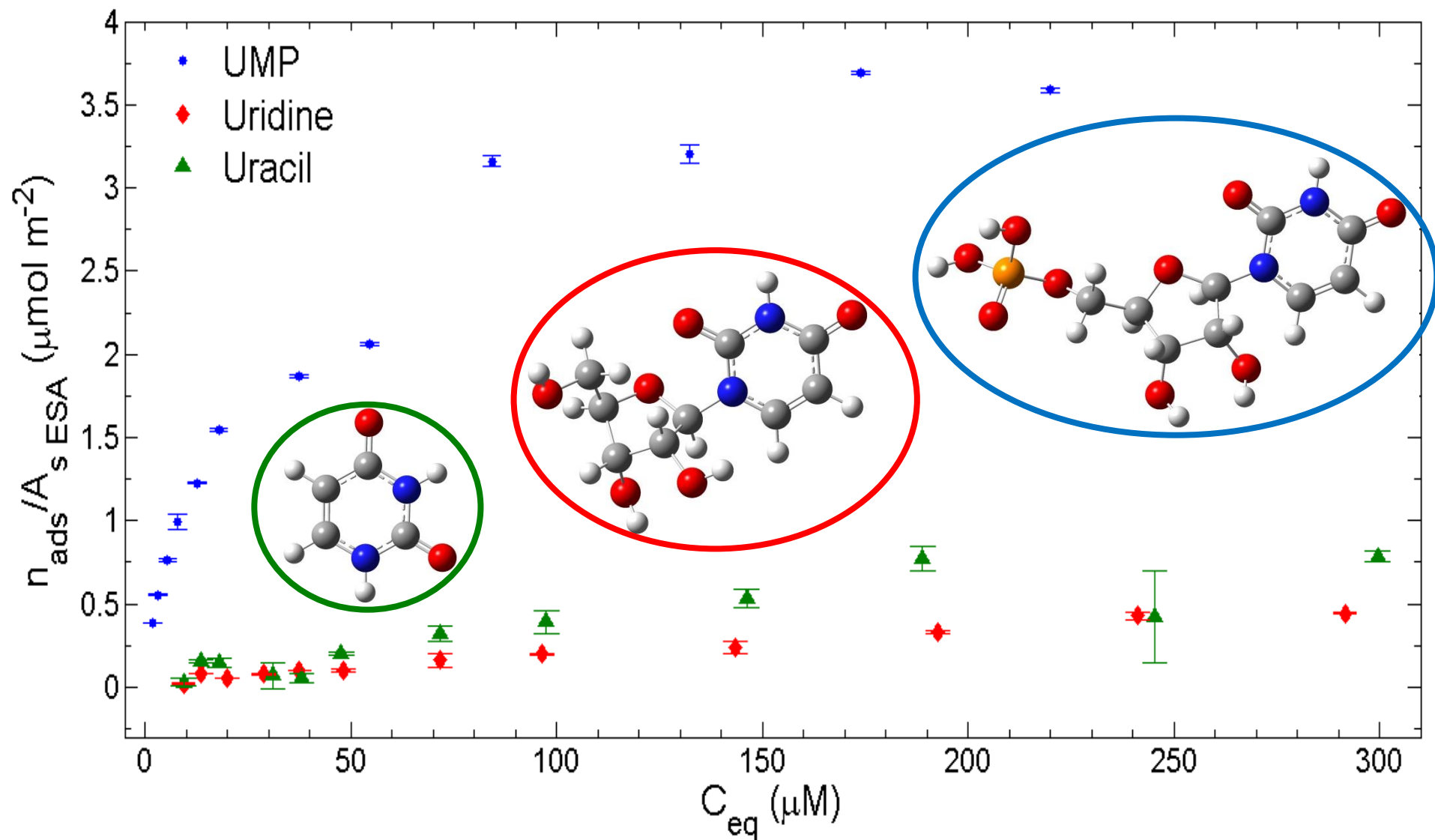


Minerals: Metal Oxides, Hydroxides and Silicates

Molecules: nucleobases, nucleosites, nucleotides, aminoacids

	DHN	Glu	Arg	Leu	Gly	Isoval	Nucleobeses	Nucleosites	Nucleotides
Oligoclasio							X	X	X
Lizardite	X				X		X	X	X
Pirite	X				X			X	X
Mimetite						X	X	X	X
Natrolite	X					X	X	X	X
Serpentinite	X				X	X	X	X	
Brucite	X				X	X	X	X	
Olivine	X				X		X	X	X
SiO2		X	X	X					

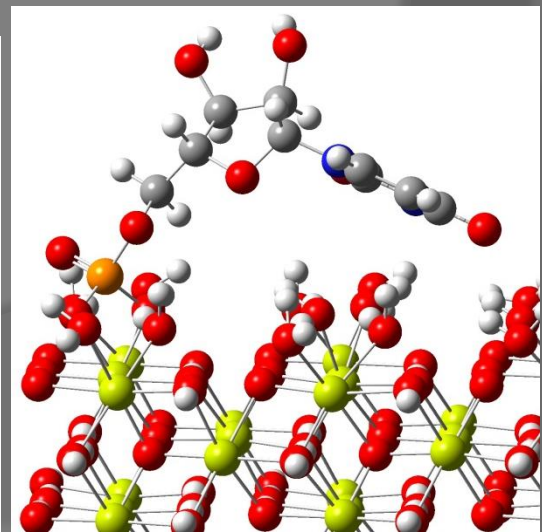
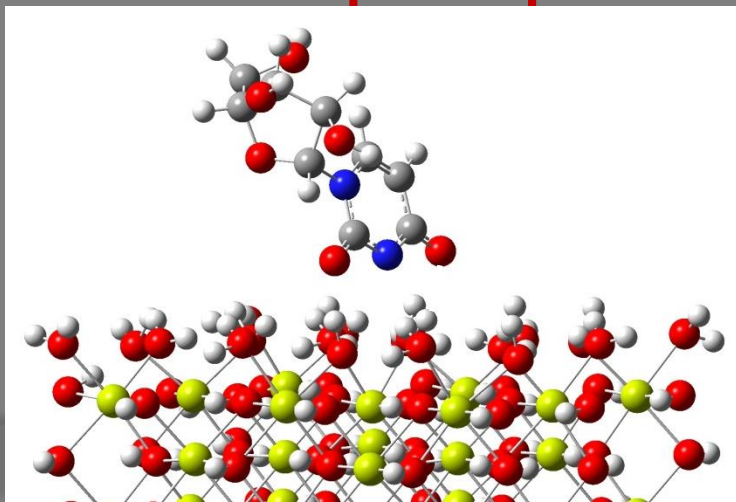
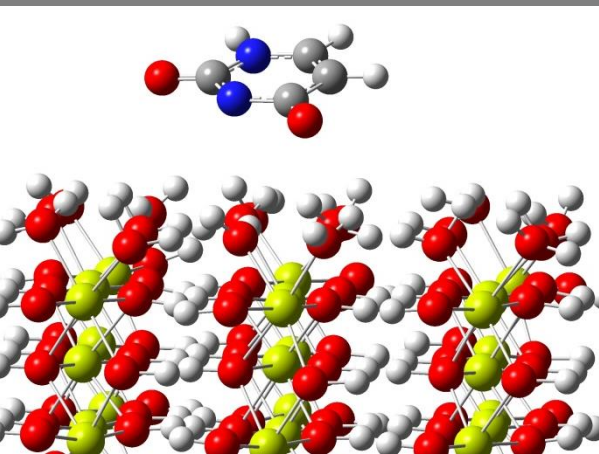
Adsorption of Uracil, Uridine and UMP on Brucite in Water



Ribose not involved in the adsorption (only weak outer-sphere interactions)
Strong interactions via Phosphate group

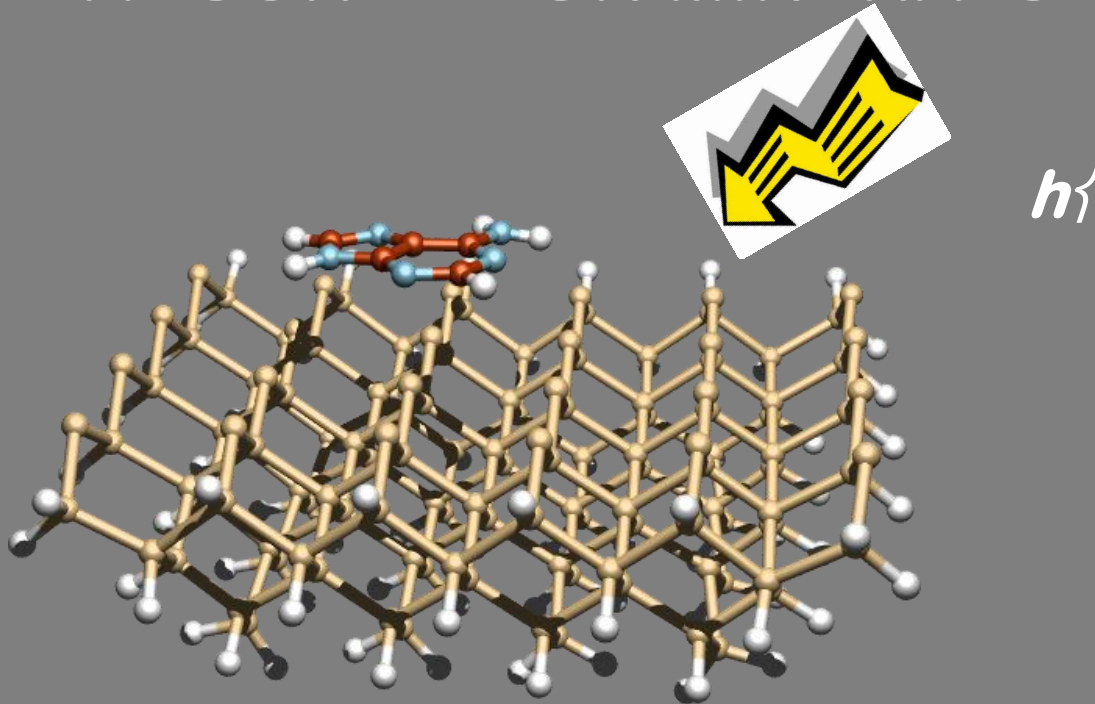
SERPENTINITE-HOSTED HYDROTHERMAL MINERALS

- Brucite selectively adsorbs nucleic acid components from dilute aqueous environments, suggesting a **role in concentrating biomolecules in prebiotic conditions**
- Brucite surface induces well-defined orientations of the molecules through specific molecule-mineral interactions, suggesting a **role in assisting prebiotic self-organization, increasing molecular complexity and promoting chemical reactions towards more complex species**

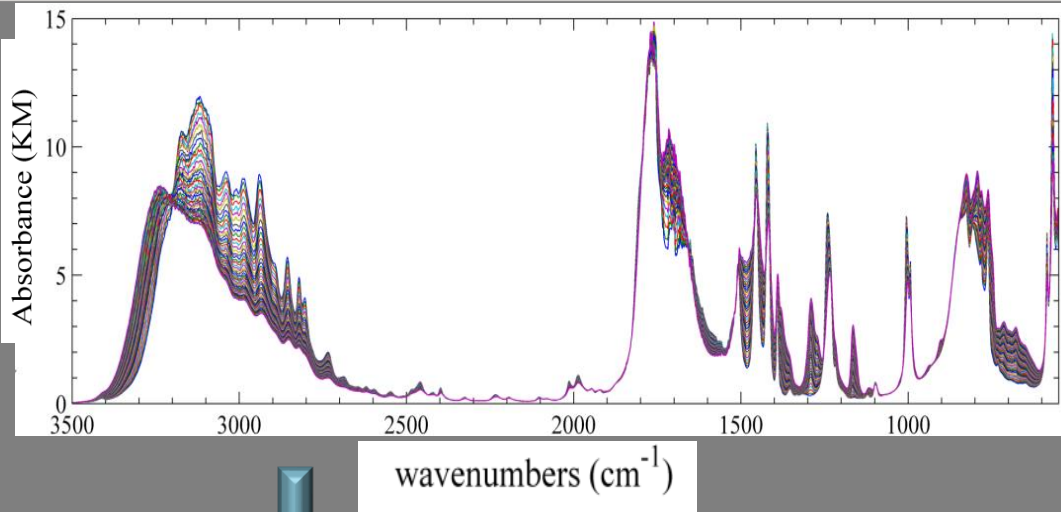


HIGH UV IRRADIATION AND SPACE-LIKE ENVIRONMENTS

***UV IRRADIATION OF “BUILDING BLOCKS OF LIFE”
ADSORBED ON MINERALS***



UV degradation kinetics



$$N(t)/N_0 = Be^{-\beta t} + c$$

$N(t)/N_0$ fraction of unaltered molecules

β degradation rate

B fraction of interacting molecules

c fraction of non-interacting molecules

$t_{1/2}$ half-lifetime

σ UV destruction cross section

Φ_{tot} total focused incident UV flux

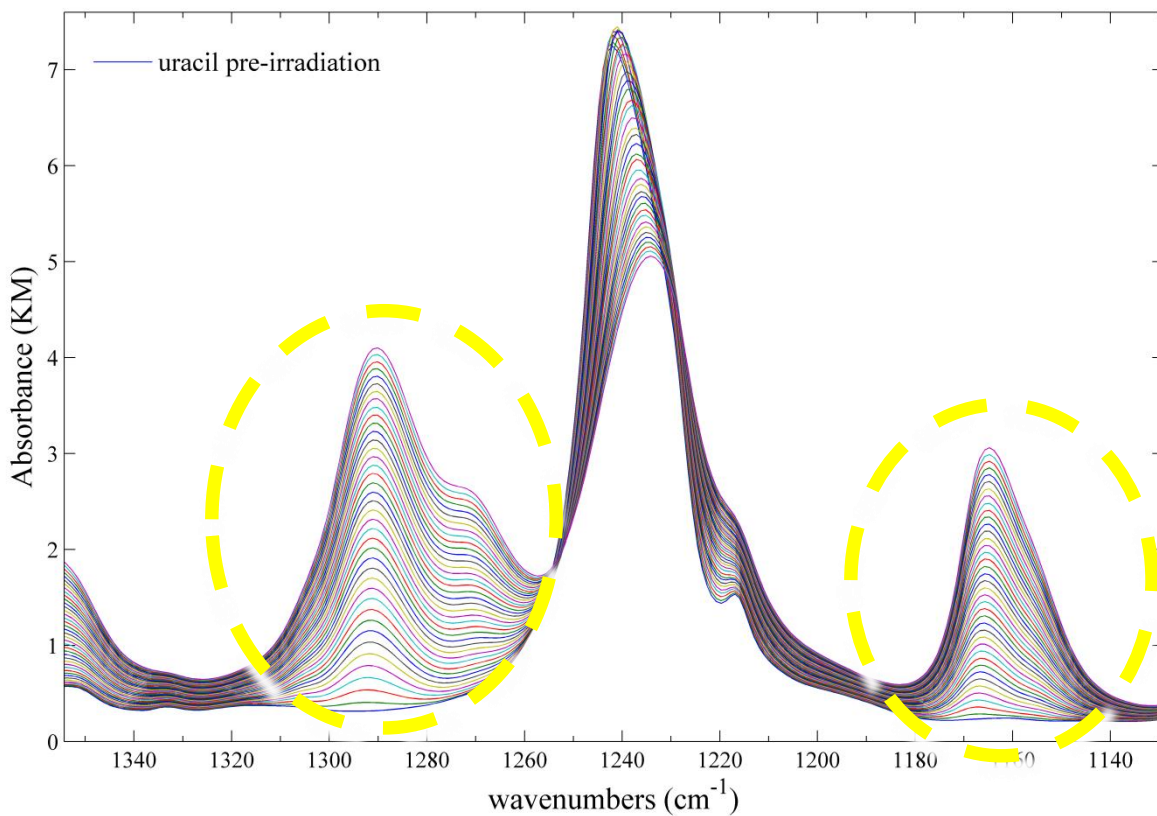
A_0 sample irradiated area

$$t_{1/2} = \ln 2 / \beta$$
$$\beta = \sigma \Phi_{\text{tot}} / A_0$$

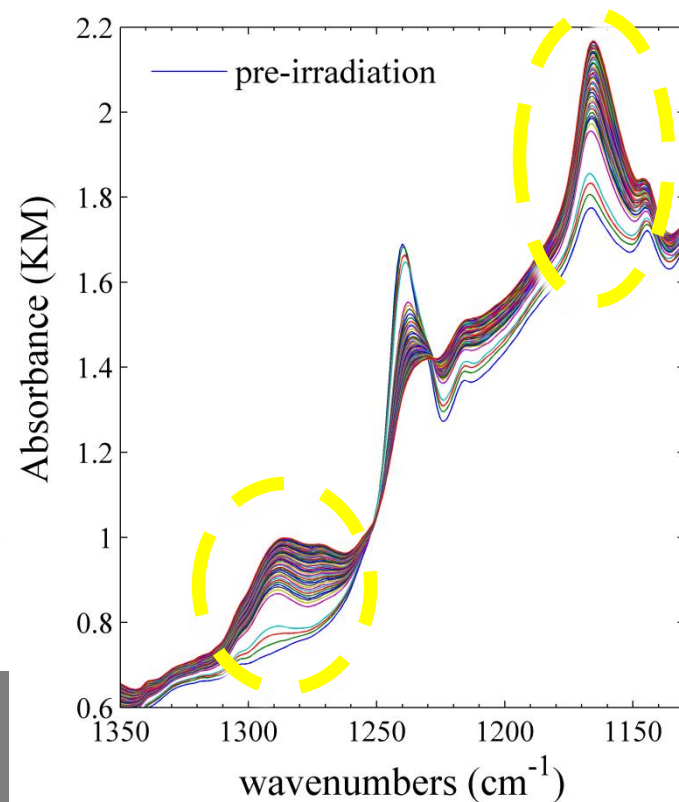
- **Cytosine** and **hypoxanthine** have a greater photostability
- For **adenine** and especially **uracil** degradation was observed both pure and adsorbed onto MgO and forsterite
- **Minerals** make degradation faster and more probable

Photoproducts marker bands

Uracil

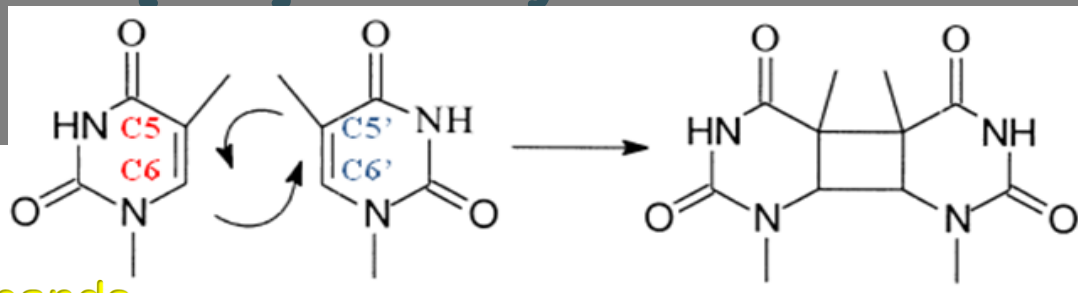


Uracil adsorbed on Forsterite

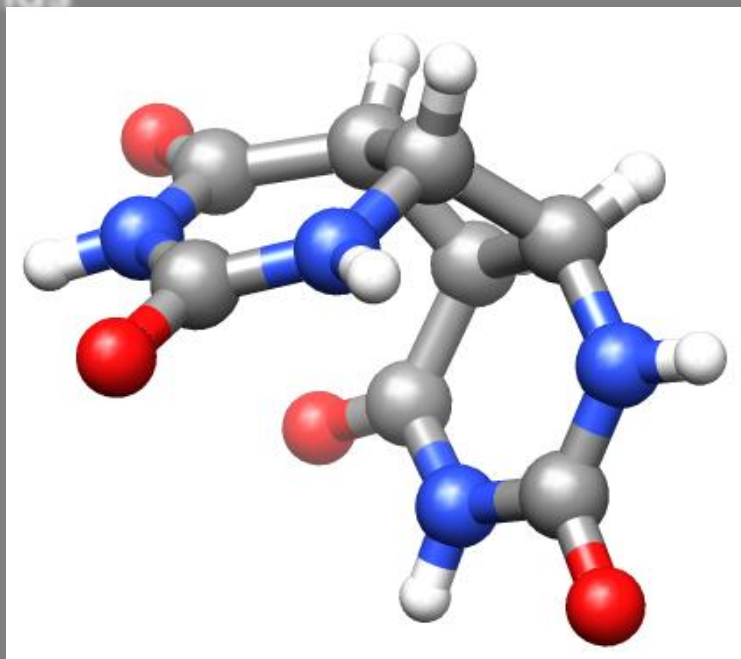


Photoproducts

[2+2] Photocycloaddition

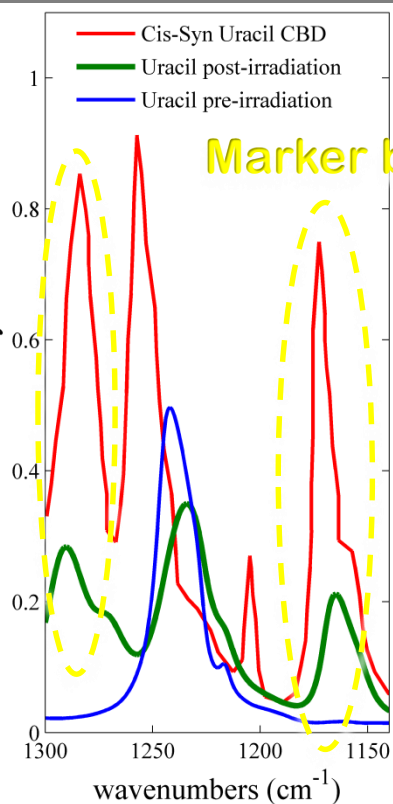


Marker bands



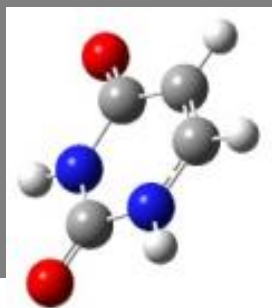
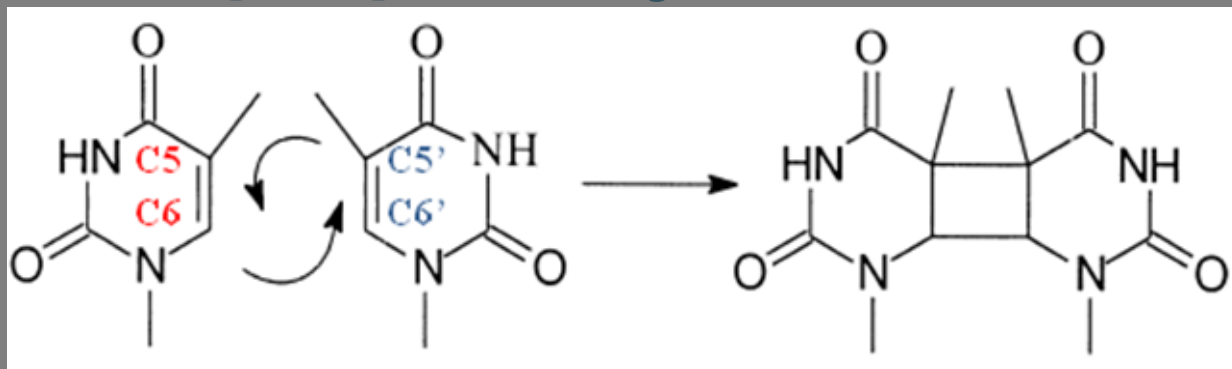
Main photoproduct:

**Cis-syn cyclobutane
dimer (CBD)**

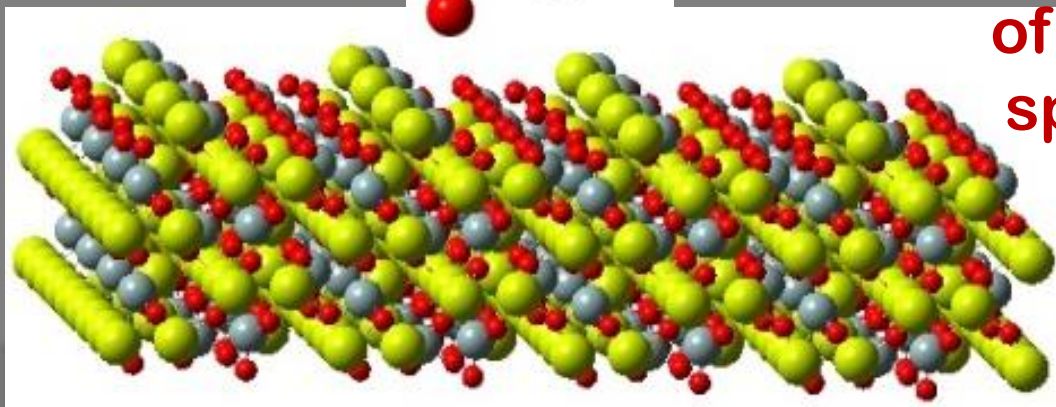


Catalytic Effect of Forsterite

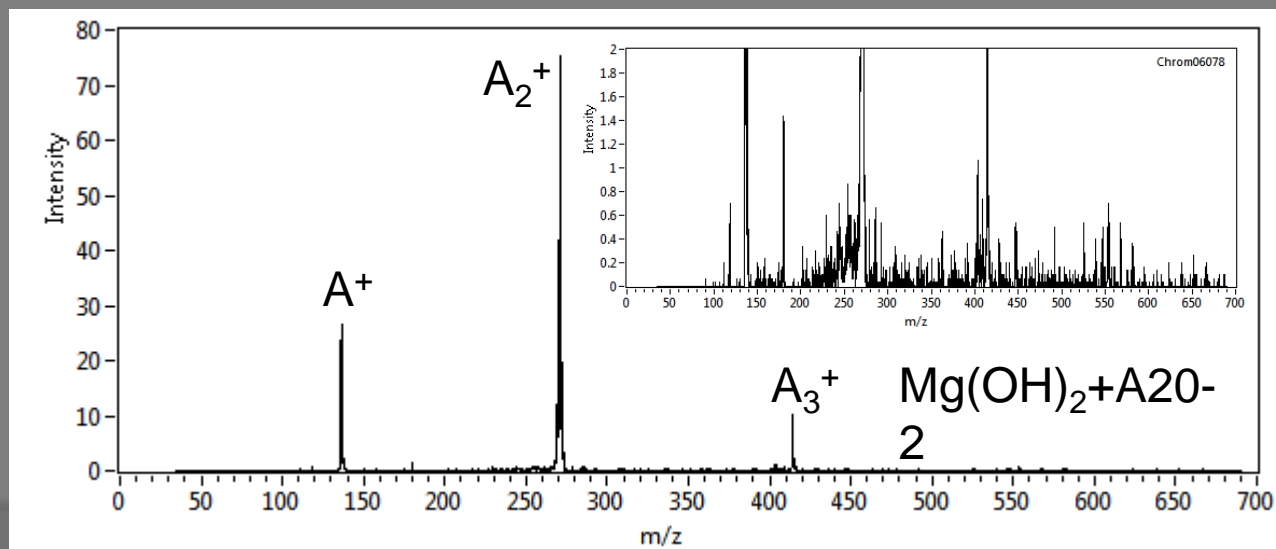
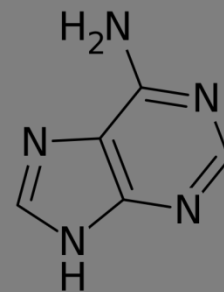
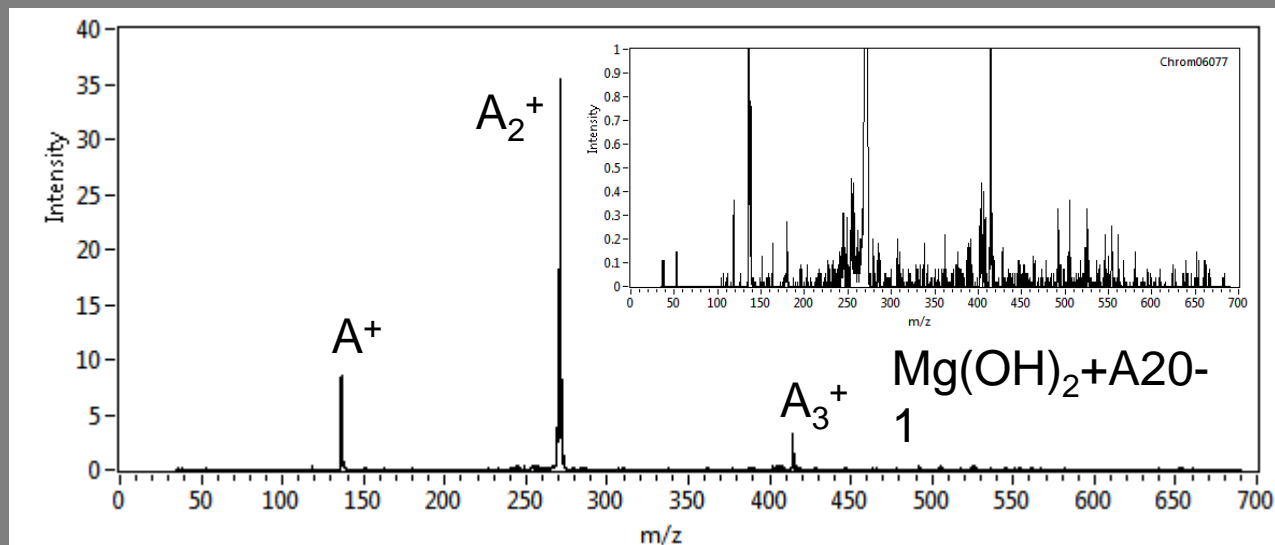
[2+2] Photocycloaddition



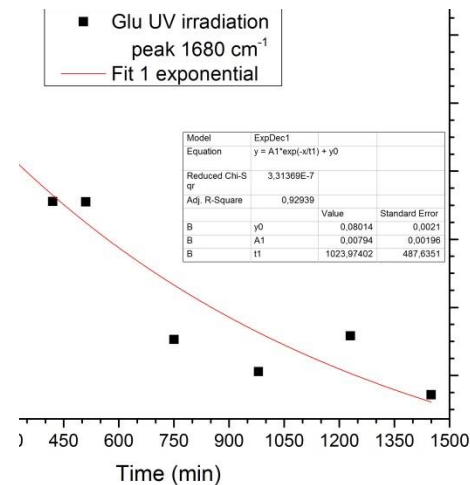
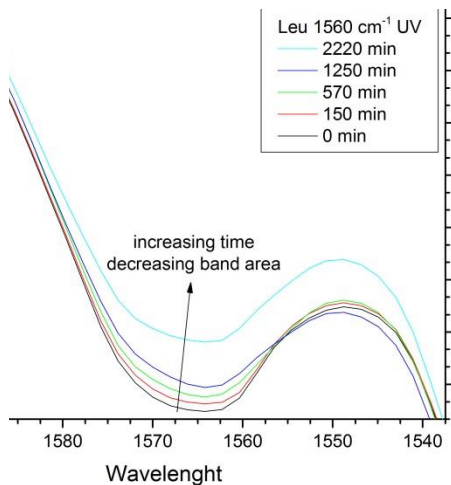
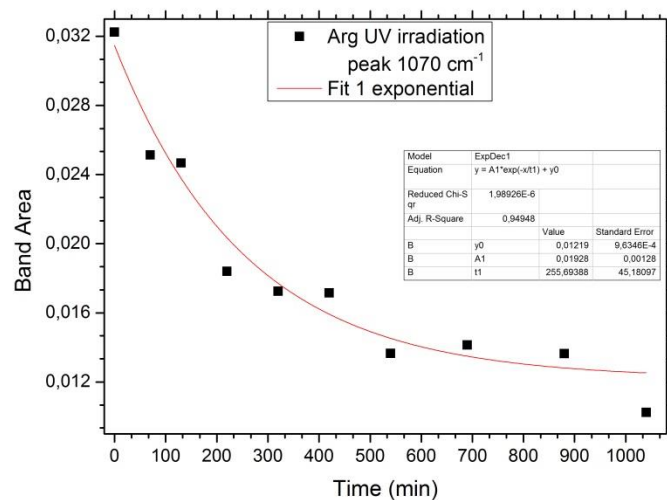
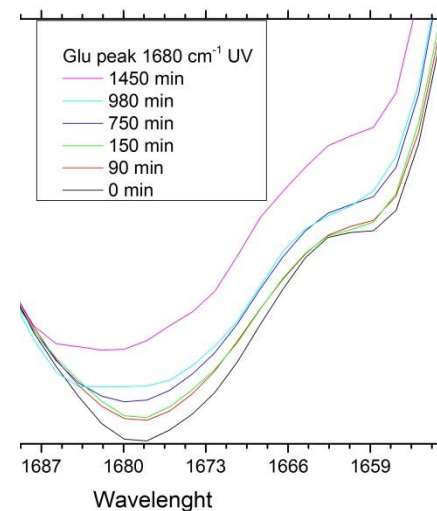
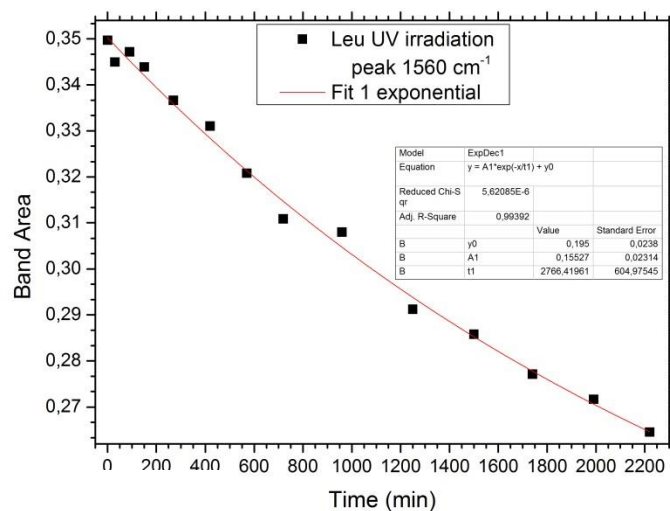
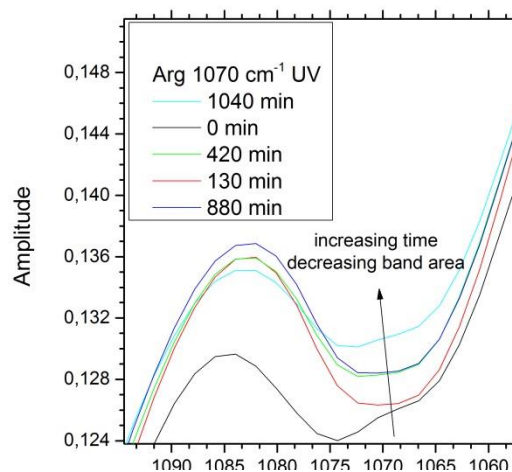
- ✓ Concentrates molecules on a local scale through adsorption
- ✓ Induces the correct orientation of reactive groups through specific molecule-mineral interactions



20 mM Adenine on Brucite



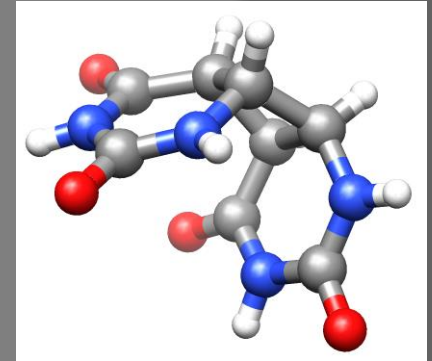
Arg and Leu UV irradiation



	Life-time (min)	Cross-section (m ²) × 10 ⁻²⁶
Glutamic acid 1680 cm ¹	1.0 ± 0.5	2 ± 1
Glutamic acid 670 cm ¹	0.6 ± 0.1	3.6 ± 0.7
Glutamic acid 1267 cm ¹	2.7 ± 3.5	0.8 ± 1.1
Leucine 1560 cm ¹	2.8 ± 0.6	0.8 ± 0.2
Leucine 670 cm ¹	2.8 ± 0.8	0.8 ± 0.3
Leucine 1530 cm ¹	2.2 ± 0.9	1.0 ± 0.5
Arginine 1070 cm ¹	0.26 ± 0.05	9 ± 2

HIGH UV IRRADIATION AND SPACE-LIKE ENVIRONMENTS

- Uracil is the most photoreactive, probably forming **cyclobutane dimers**
- **MgO** and **Forsterite** have no protective effect, instead they **may be catalytic potentially triggering chemical processes towards more complex species**



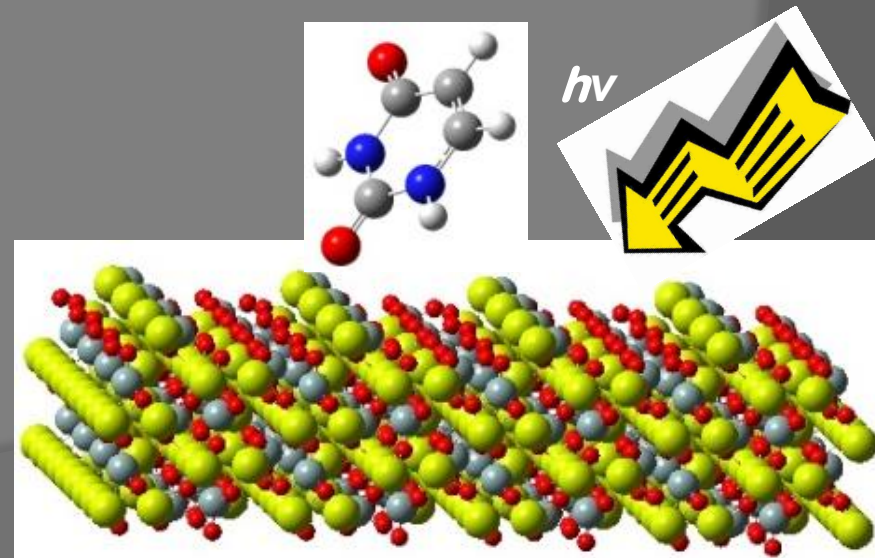
Fornaro, T.; Brucato, J. R.; Pace, E.; Guidi, M. C.; Branciamore, S.; Pucci, A. *Icarus* **2013**, 226, 1068-1085.

Fornaro T.; Brucato, J. R.; Branciamore, S.; Pucci, A. *International Journal of Astrobiology* **2013**, 12 (1), 78-86.

TAKE-HOME MESSAGES...

Mineral surfaces have the ability **to selectively adsorb** and **concentrate organic molecules** on a local scale, removing key organics from aqueous environments, **promoting self-organization** through specific molecule-mineral interactions

Minerals can act as **photocatalysts** promoting reactions towards more complex species



Acknowledgements

❖ **INAF- Astrophysical Observatory of Arcetri (Florence):** **Dr. Teresa Fornaro.**



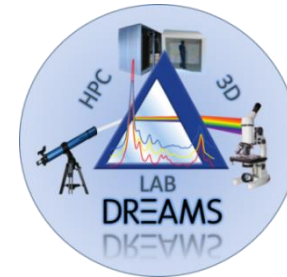
❖ **Italian Space Agency (ASI):**
Grant I/060/10/0-ExoMars Science.



❖ **Laboratori Nazionali di Frascati (LNF) – INFN:**
Dafne Luce Laboratory.



❖ **Scuola Normale Superiore di Pisa:**
Prof. Vincenzo Barone, Dr. Malgorzata Biczysko, DREAMS team.



Dedicated Research Environment for Advanced Modeling and Simulation

❖ **Geophysical Laboratory - Carnegie Institution for Science, Washington DC, USA:**

Prof. Robert Hazen, Dr. Cecile Feuillie,

Prof. Dimitri Sverjensky (Johns Hopkins University).

